Light Source I

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- Introduction
 Fundamentals of Light and SR
 - Overview of SR Light Source
 Characteristics of SR (1)



- Light Source II
 Characteristics of SR (2)
 Practical Knowledge on SR

Introduction

SR Facility and Light Source

- SR: Definition
 - Electromagnetic wave emitted by a charged particle deflected by a magnetic force
- SR Facility
 - Accelerators to generate a high-energy electron beam
 - Magnetic devices (SR light source) to generate intense SR
 - Optical elements (monochromators, mirrors,..)
 - Experimental stations

SR as a Probe for Research

- SR has a lot of advantages over other conventional light sources
 - Highly collimated (laser-like)
 - Wavelength tunability
 - Polarization
 - **—**
- However, the total radiation power does not differ significantly.

Comprehensive understanding of SR (and light source) is required for efficient experiments.

Topics in This Lecture (1)

- Fundamentals of Light and SR
 - Why we need SR?
 - Physical quantity of light
 - Uncertainty of light: Fourier and diffraction limits
 - SR: Light from a moving electron
- Overview of SR Light Source
 - Types of light sources
 - Magnet configuration
- Characteristics of SR (1)
 - Radiation from bending magnets

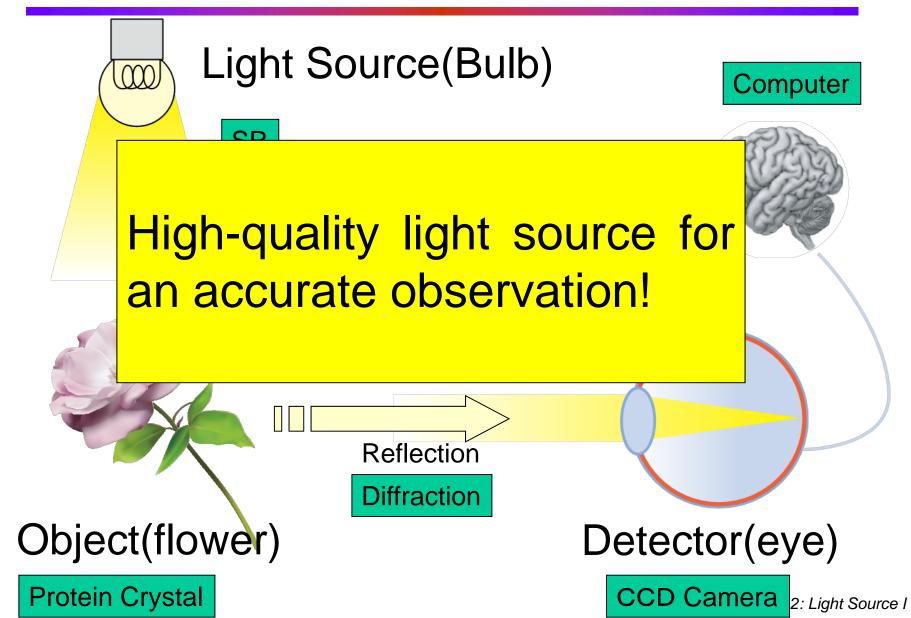
Topics in This Lecture (2)

- Characteristics of SR (2)
 - Electron Trajectory in IDs
 - Radiation from wigglers
 - Radiation from undulators
- Practical Knowledge on SR
 - Finite emittance and energy spread
 - Heat load and photon flux
 - Evaluation of optical properties of SR
 - Definition of undulators and wigglers
 - Numerical examples

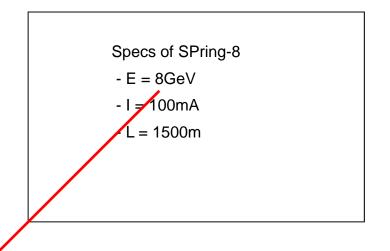
Fundamentals of Light and SR

- Why we need SR?
- Physical Quantity of Light
- Uncertainty of Light
- SR: Light from a Moving Electron

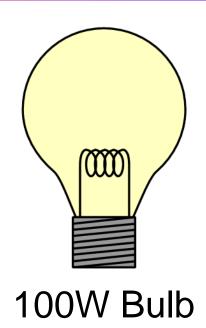
Observation with Light



Which Quality is Better?



1mW Laser (pointer)



Bulb

Lighting equipment in a room: Pointer during a presentation: Laser

Depends on the Object!

How to Define the Quality of Light?(1)

- The performance of the light source depends on the dimension of the object and the method to detect light.
- For observation, the photons emitted by the light source should be
 - illuminated on the object for interaction
 - recognized by the detector for analysis

Quality of Light Source: How efficiently the above conditions are satisfied?

How to Define the Quality of Light?(2)

Important Features of the Light Source

	Object		Related Items
	Flower	Protein	
Radiation Power	0	0	# Emitted Photons
Source Size	×	<u></u>	Illuminated Area
Directivity	Δ	0	
Monochromaticity	Δ		Accuracy of Analysis



What is Brilliance?

Brilliance(photons/sec/mm²/mrad²/0.1%B.W.)
Total Power

Source Size x Angular Divergence x Band Width

- Brilliance specifies the quality of light for observation of microscopic objects.
- The brilliance of a light source with a high total power is not necessarily high.

Examples of Brilliance

	Bulb	Laser Pointer
Total Power (W)	100	10-3
Angular Div. (mrad²)	$4\pi x 10^{6}$	1
Source Size: (mm ²)	10 ²	1
Bandwidth: (%)	100	0.01
Brilliance	~108	~10 ¹⁶
(photons/sec/)		

Laser is the best light source to observe the microscopic object!

X ray as a Probe

- Definition (not unique)
 - Electromagnetic wave (= light) with λ of 10 nm(10⁻⁸ m) ~ 0.1 Å (10⁻¹¹ m)
- Properties
 - High Energy/Photon
 - High Penetration (Roentgen etc..)
- Application to Microscopic Objects
 - X-ray Diffraction
 - Fluorescent X-ray Analysis
- No Practical Lasers!!

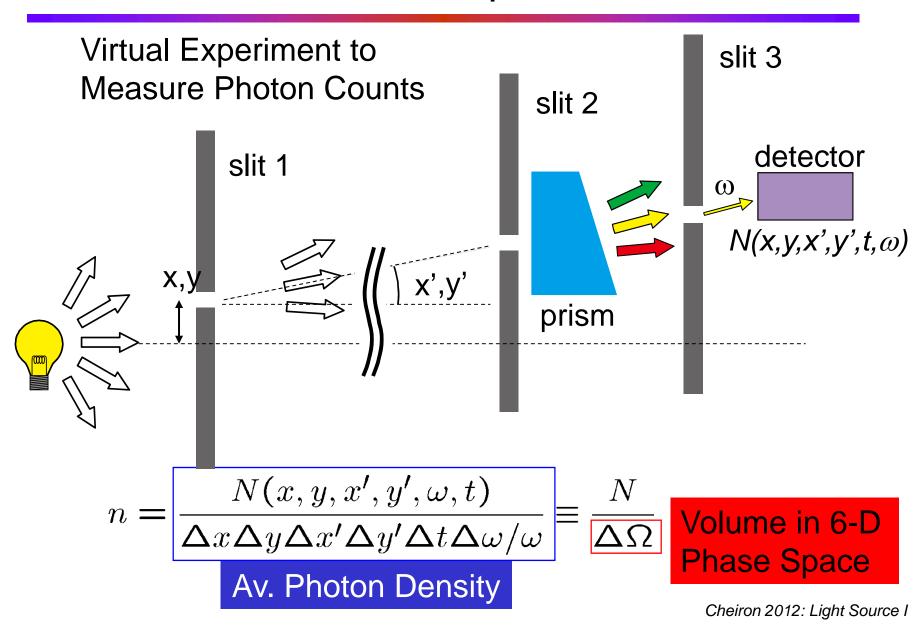


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Fundamentals of Light and SR

- Why we need SR?
- Physical Quantity of Light
- Uncertainty of Light
- SR: Light from a Moving Electron

Phase Space



Brilliance (Brightness)

Brilliance (photons/sec/mm²/mrad²/0.1%B.W.)
is defined as the photon density in the 6D
phase space, i.e.,

$$B = \lim_{\Delta\Omega \to 0} n = \frac{d^{6}N(x, y, x', y', t, \omega)}{dxdydx'dy'dtd\omega/\omega}$$

 In practice, ΔΩ can never be 0 due to uncertainty of light, thus brilliance is not a physical quantity that can be actually measured.

Photon Flux and Flux Density

 Removing the 1st slit gives the angular flux density (photons/sec/mrad²/0.1%B.W), i.e.,

$$\frac{d^2F}{dx'dy'} = \iint Bdxdy$$

 Removing the 1st & 2nd slits gives the total flux (photons/sec/0.1%B.W), i.e.,

$$F = \iiint B dx dy dx' dy'$$

 Estimation of number of photons to be delivered to the sample.

Radiation Power and Power Density

 Removing the 1st & 3rd slits gives the angular power density (W/mrad²), i.e.,

$$\frac{d^2P}{dx'dy'} = 10^3 Q_e \pi \int \frac{d^2F}{dx'dy'} d\omega$$
conversion from photons/sec/0.1%B.W. to W

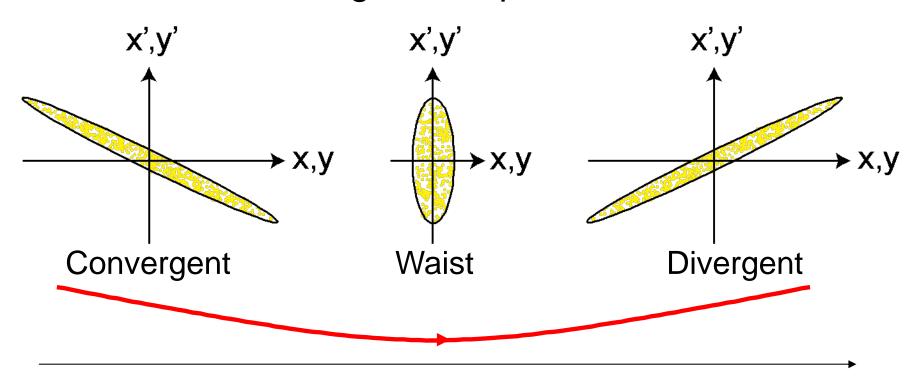
 Removing all the slits gives the total power (W), i.e.,

$$P = 10^3 Q_e \hbar \iiint \frac{d^2 F}{dx' dy'} d\omega dx' dy'$$

Estimation of heat load on BL components.

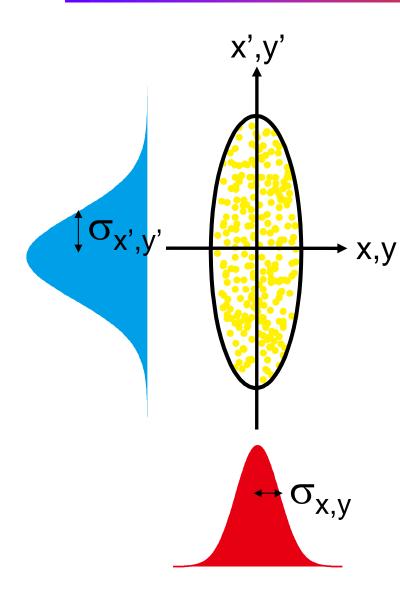
Photons in 4D Phase Space

 Photon distribution in the 4-D phase space at different longitudinal positions.



Beam Envelope along Propagation Axis

Source Size, Divergence, Emittance



- Source size $(\sigma_{x,y})$ is defined as the beam envelope at the beam waist position.
- Angular divergence $(\sigma_{x',y'})$ is constant along the axis of propagation, as far as no optical elements are present.
- Emittance $(\varepsilon_x, \varepsilon_y)$ is defined as $\sigma_{x,y} \times \sigma_{x',y'}$, which is equal to the area of the phase ellipse divided by π .

Fundamentals of Light and SR

- Why we need SR?
- Physical Quantity of Light
- Uncertainty of Light
- SR: Light from a Moving Electron

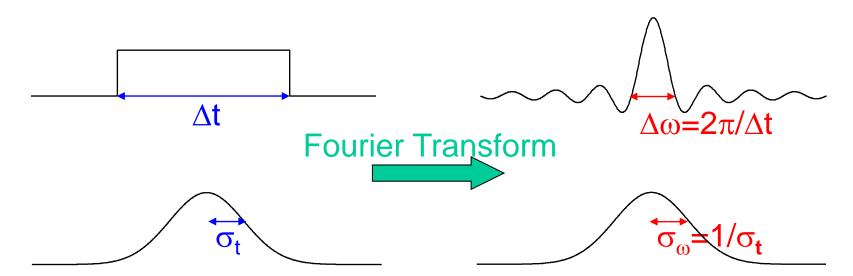
Uncertainty of Light

- The photon distribution in the 6D phase space (x,y,x',y',t,ω) gives us the full information on the properties of SR.
- Due to wave nature of light, however, we have two uncertainty relations to take care, which are well characterized by the Fourier transform.
- These relations impose two restrictions on SR, Fourier and Diffraction limits.

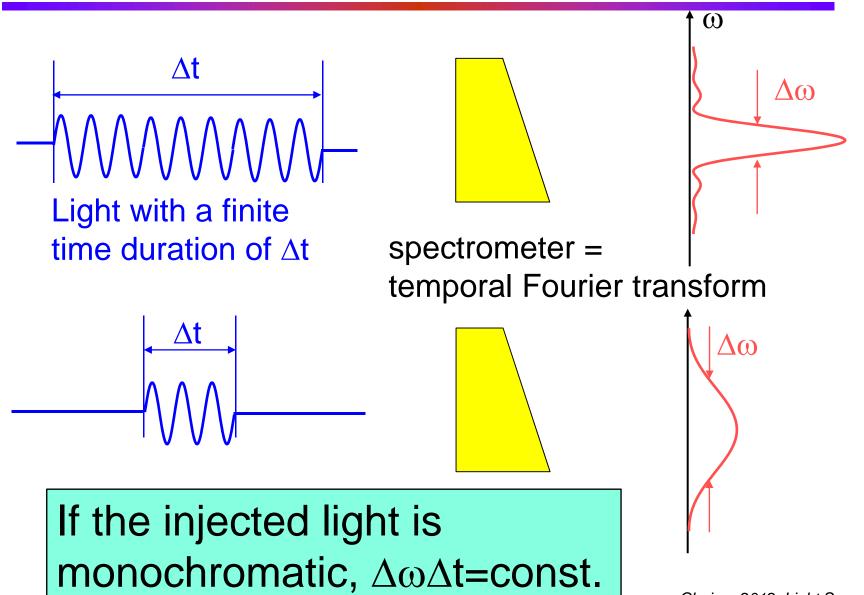
Fourier Transform: Example

Important Fourier Transform in SR Formulae

F(t)	$f(\omega) = \int_{-\infty}^{\infty} F(t) e^{i\omega t}$
$\begin{cases} 1/\Delta t; & -\Delta t/2 \le t \le \Delta t/2 \\ 0; & t < -\Delta t/2, \Delta t/2 < t \end{cases}$	$\frac{\sin \omega \Delta t/2}{\omega \Delta t/2} \equiv \operatorname{sinc}(\omega \Delta t/2)$
$\frac{1}{\sqrt{2\pi}\sigma_t}\exp\left(-\frac{t^2}{2\sigma_t^2}\right)$	$\exp\left(-\frac{\omega^2\sigma_t^2}{2}\right)$



Spectrum of Light



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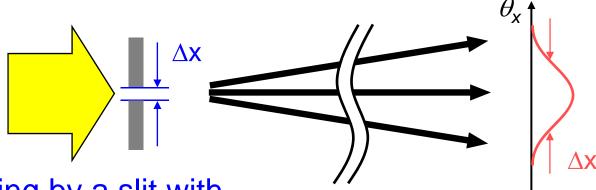
Fourier Limit of Light

Temporal Fourier transform imposes

$$\Delta\omega\Delta t \geq \text{const.}$$

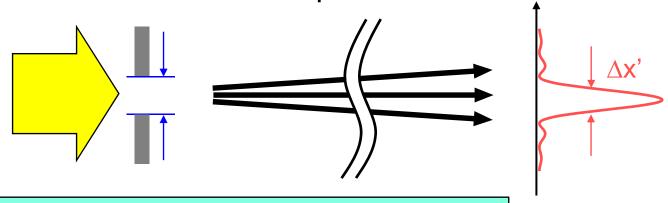
- Uncertainty of light in the (ω,t) plane.
- When equality holds, light is said to be
 - Fourier-limited
 - Temporally coherent
- Important to understand the spectral properties of SR.

Diffraction Pattern of Light



Clipping by a slit with a finite width of Δx

diffraction pattern in the far region = spatial Fourier transform



If the injected light is a plane wave, then $\Delta x \Delta x' = const.$

Diffraction Limit of Light

Spatial Fourier transform imposes

$$\Delta x \Delta k_x \ge \text{const.}$$

$$k_{x}=(2\pi/\lambda)x'$$
 $\Delta x \Delta x' \geq \lambda \times \text{const.}$

- Uncertainty of light in (x,x') plane
- When equality holds, light is said to be
 - Diffraction limited
 - Spatially coherent
- In the case of Gaussian beam,

$$\sigma_x \sigma_{x'} \geq \lambda/(4\pi)$$
 Natural emittance of light

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Fundamentals of Light and SR

- Why we need SR?
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- SR: Light from a Moving Electron

SR: Light from a Moving Electron

- Unlike the ordinary light source (sun, light bulb,...), the light emitter of SR (electron) is ultra-relativistic.
- The characteristics of SR is thus quite different because of relativistic effects.
- What we have to take care is:
 - 1. Speed-of-light limit
 - 2. Squeezing of light pulse
 - 3. Conversion of the emission angles

Speed-of-Light Limit

Within the framework of relativity, the velocity of an *electron* never exceeds the speed of light.

$$v/c = \beta = \sqrt{1 - \gamma^{-2}}$$

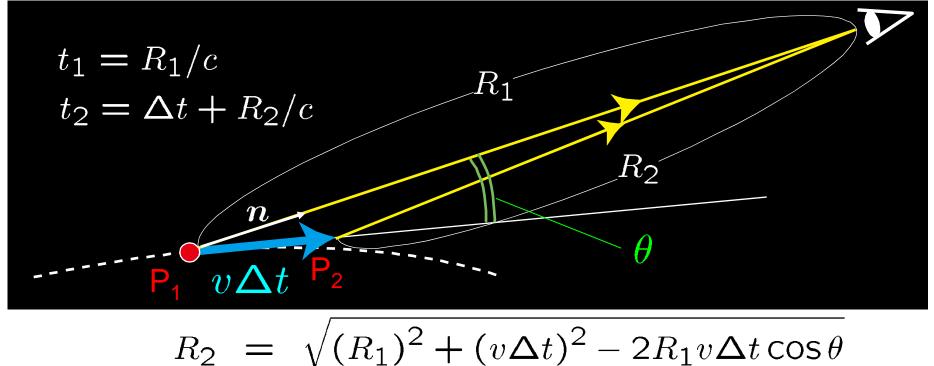
$$\sim 1 - \frac{1}{2\gamma^2}$$

Energy	β
1MeV	0.941
10MeV	0.9988
100MeV	0.999987
8GeV	0.99999998

$$\gamma = \frac{E}{mc^2}$$

:Lorentz Factor (relative electron energy,mc²=0.511MeV)

Squeezing of Light Pulse Duration



$$R_2 = \sqrt{(R_1)^2 + (v\Delta t)^2 - 2R_1 v\Delta t \cos \theta}$$

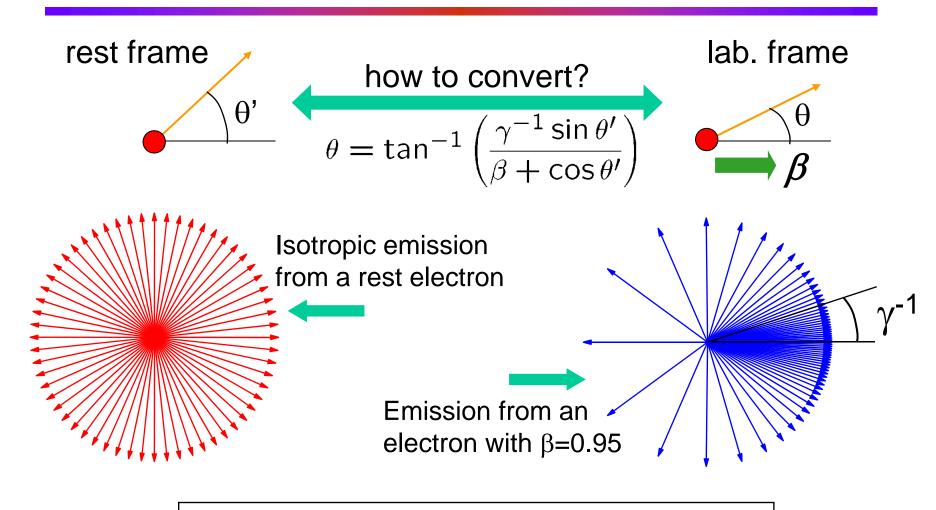
$$\sim R_1 - (v \cdot n)\Delta t$$

$$\Delta \tau = t_2 - t_1 = \Delta t + R_2/c - R_1/c$$

$$= \Delta t \left(1 - \beta \cdot n\right) = \frac{\Delta t}{2\gamma^2} \gamma >> 1, \theta = 0$$
time squeezing

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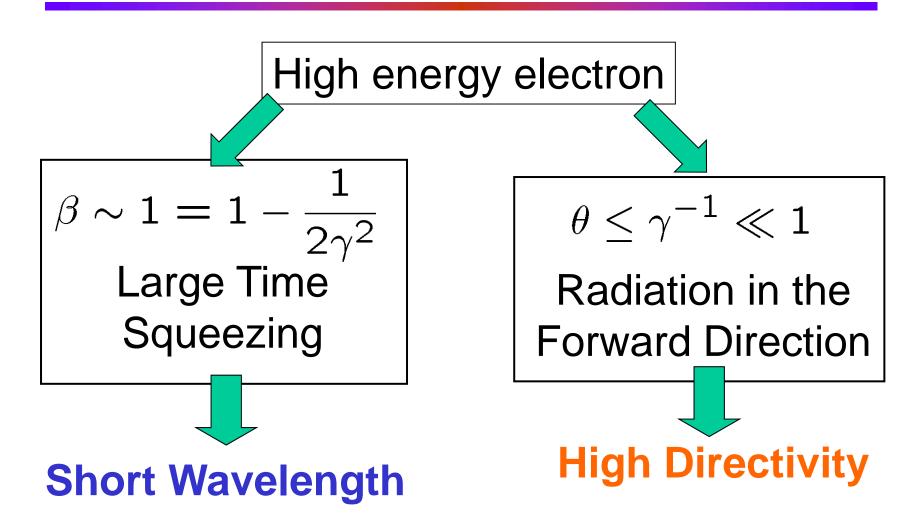
Conversion of Emission Angles



Light emitted from a moving object $(\beta\sim1)$ concentrates within γ^{-1}

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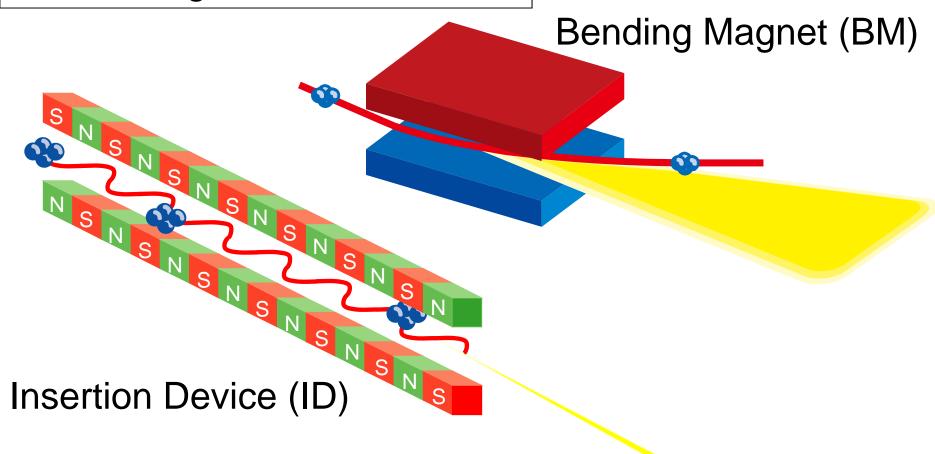
SR from a High-Energy Electron



Overview of SR Light Source

What is SR Light Source?

Magnets to deflect the electron beam and generate SR.



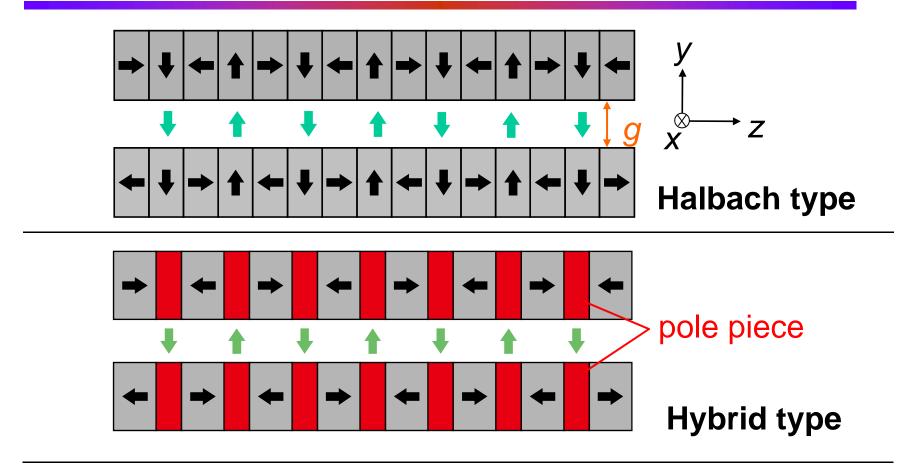
Bending Magnet

- One of the accelerator components in the storage ring.
- Generate uniform field to guide the electron beam into a circular orbit.
- EMs combined with highly-stable power supplies are adopted in most BMs due to stringent requirement on field quality and stability.
- Superconducting magnets are used in a few facilities in pursuit of harder x rays.

Insertion Device

- Installed (inserted) into the straight section of the storage ring between two adjacent BMs.
- Generate a periodic magnetic field to let the injected electron beam move along a periodic trajectory.
- Most IDs are composed of PMs, while EMs are used for special use such as helicity switching.
- Classified into wigglers and undulators.

Magnetic Circuit of IDs

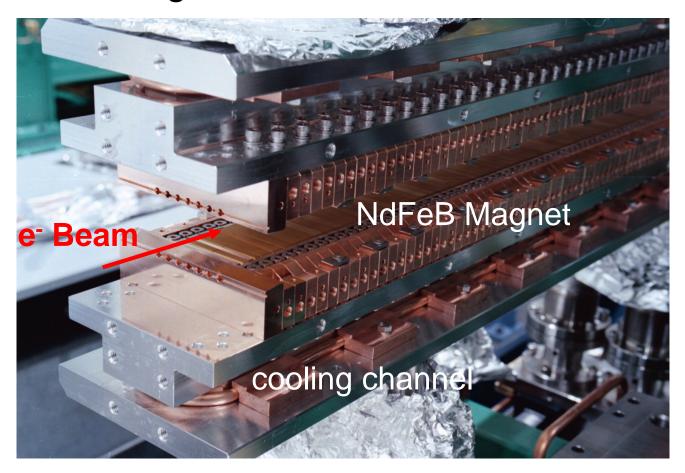


In each type, a sinusoidal magnetic field is obtained:

$$B_y(z) \sim B_0(B_r, g/\lambda_u) \sin\left(\frac{2\pi z}{\lambda_u}\right)$$

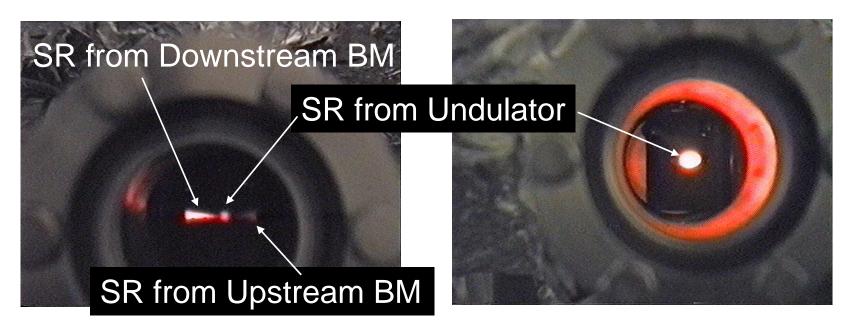
Example of ID Magnets

Halbach-type Magnet Array for SPring-8 Standard Undulators



Example of SR Image

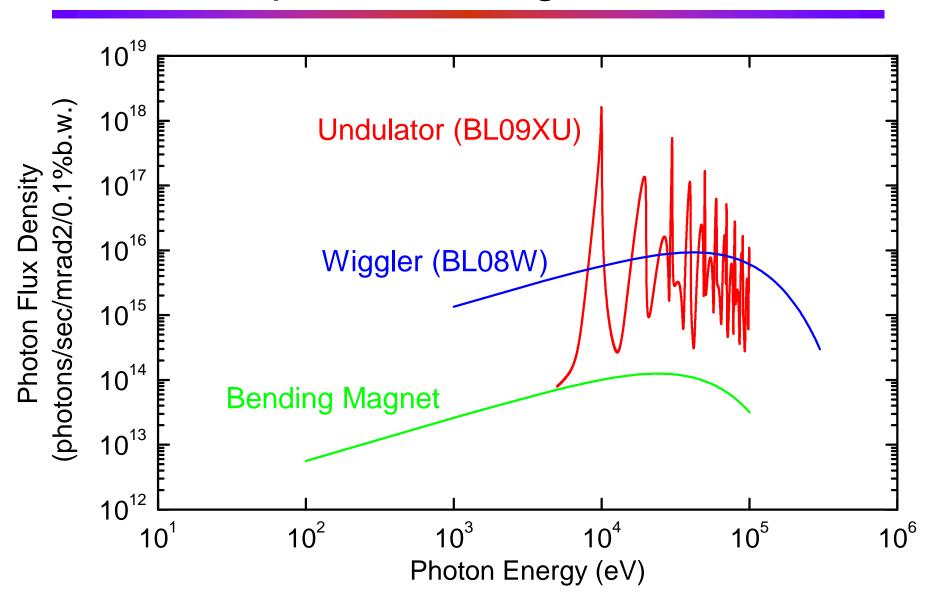
BL41XU@SP-8, first image of SR with a fluorescent screen (<0.1mA)



Undulator Gap = 50 mm

Undulator Gap = 20 mm

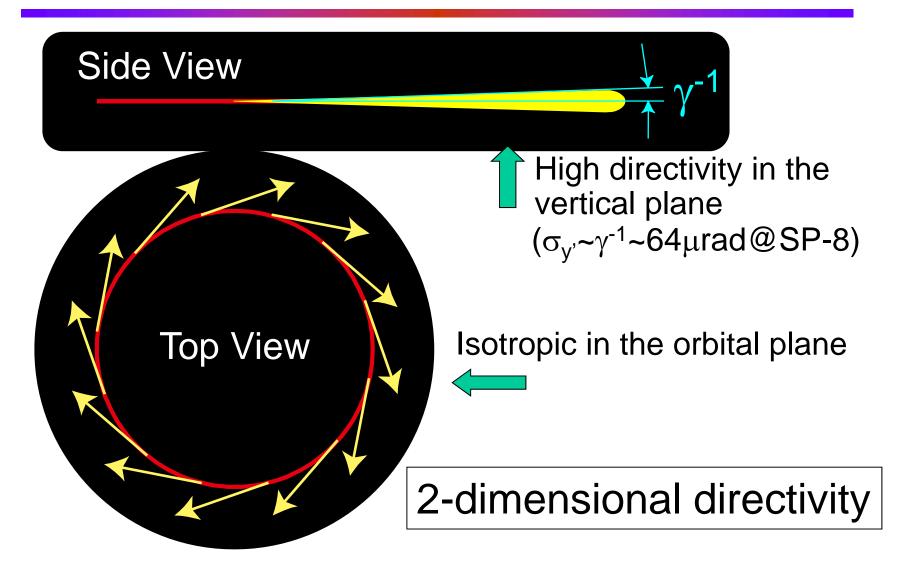
Comparison of Light Sources



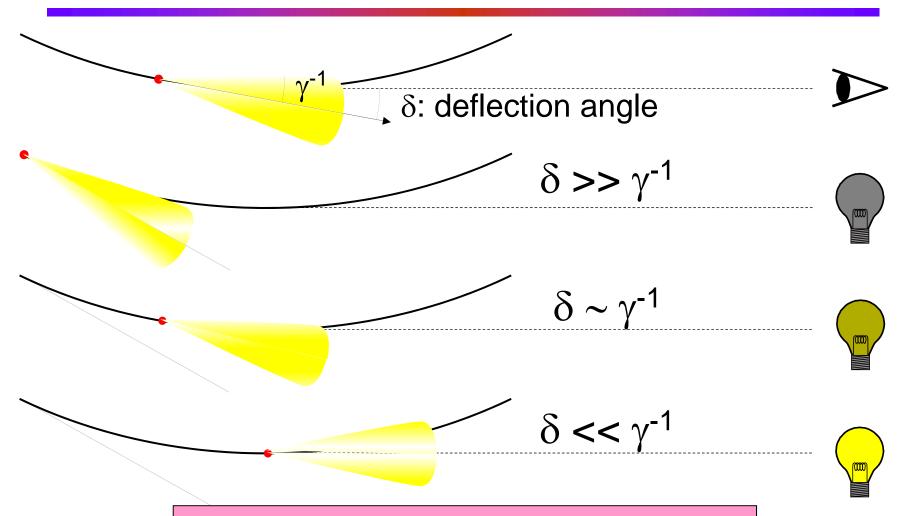
Characteristics of SR (1)

Radiation from BMs

Directivity of BM Radiation

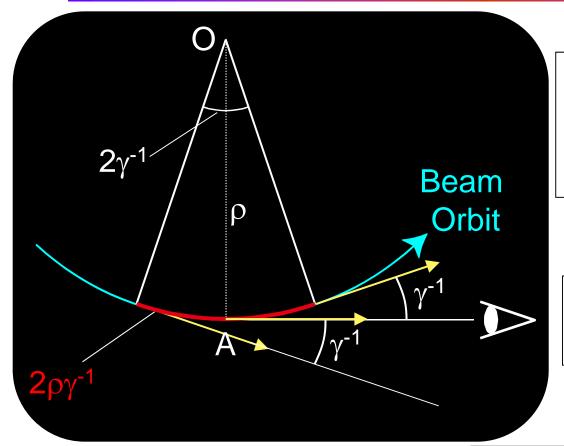


Spectrum of BM Radiation (1)



Photons emitted when $\delta < \gamma^{-1}$ are detected by the observer

Spectrum of BM Radiation (2)

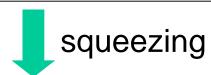


Major contribution of radiation is from the portion painted red



Pulse duration for e

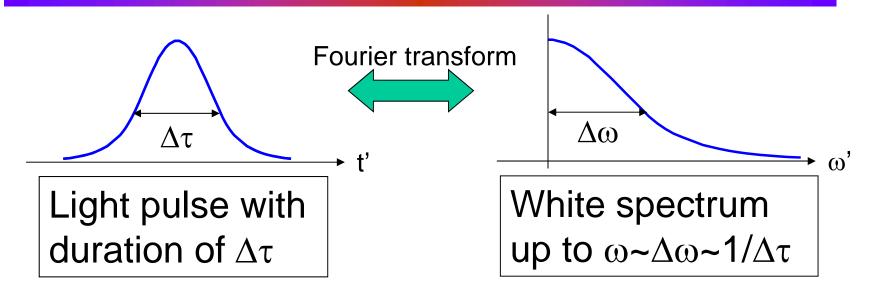
$$\Delta t = 2\rho \gamma^{-1}/c$$



Pulse duration for observer

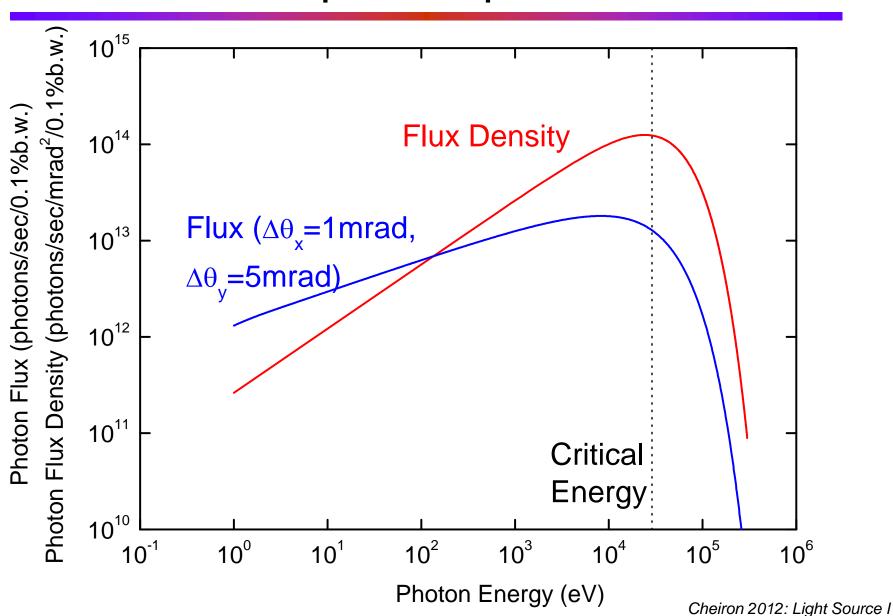
$$\Delta \tau = \frac{\Delta t}{2\gamma^2} = \frac{\rho}{\gamma^3}$$

Spectrum of BM Radiation (3)

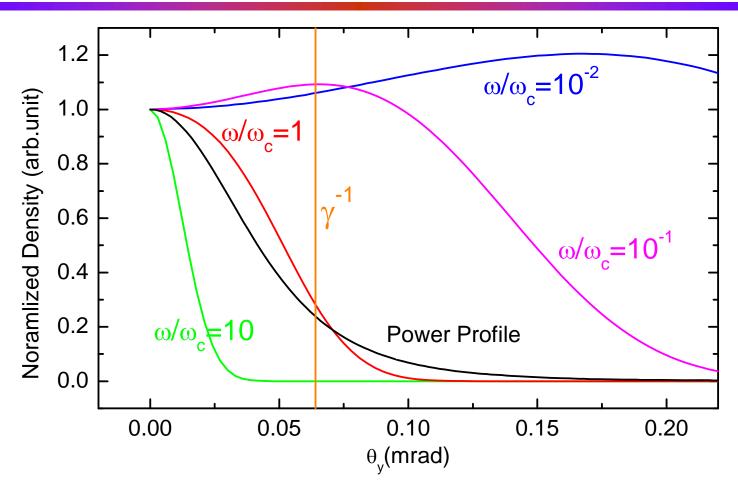


- By definition, $\omega_c = (3/2)/\Delta \tau = 3\gamma^3 c/2\rho$ is called "critical frequency" of SR, which gives a criterion of the maximum energy of SR from a BM.
- In practical units,
 ħω_c(keV)=0.665E_e²(GeV)B(T)

Example of Spectrum

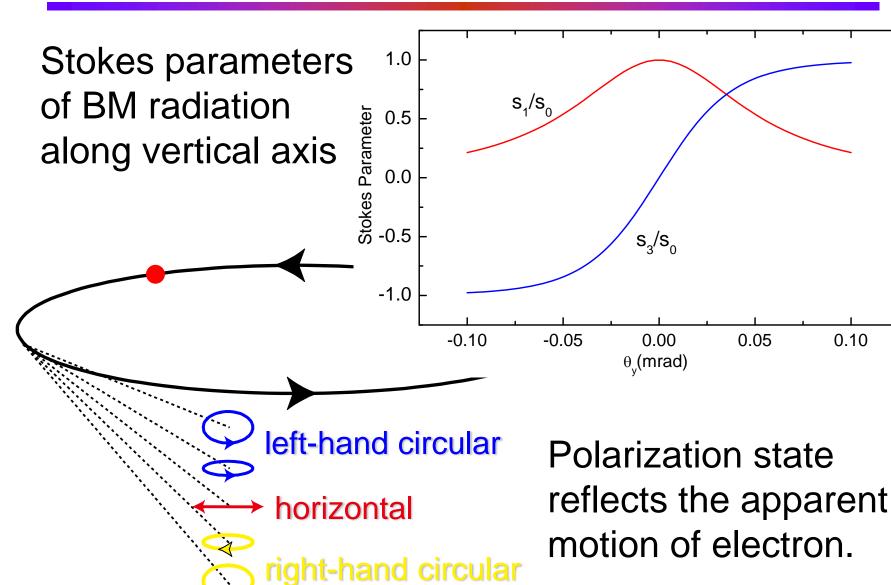


Angular Profile of BM Radiation



- power profile ~ flux profile $@\omega/\omega_c=1$
- larger angular divergence for lower energy

Polarization of BM Radiation



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